

Complexity of Quantum Field Theories (QFTs)

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Quantum Complexity Theory 6.845
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Outline

- 1 Introduction and Motivation
- 2 Intuition and Complications
- 3 Algorithm Overview



Outline

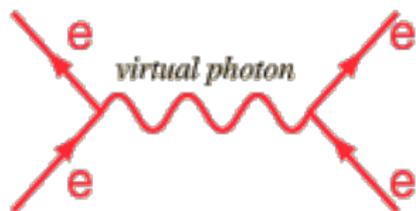
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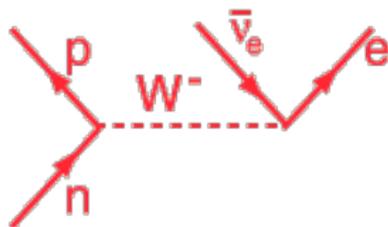
QFT: Marriage of QM and Special Relativity

Special Relativity: $E = m(c^2)$.
Infinitely Many Degrees of Freedom
 \Rightarrow QFT

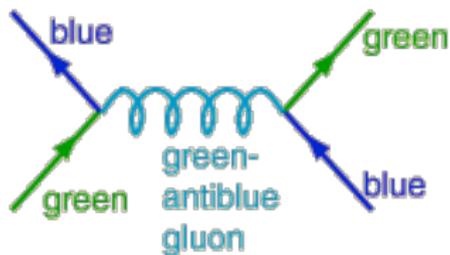
Standard Model



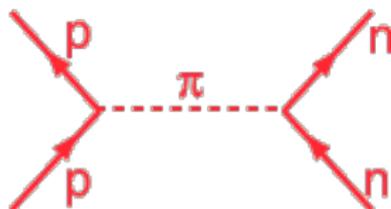
Electromagnetic



Weak



between quarks



between nucleons

Strong Interaction

Describing a QFT

Many different pictures of QM.

- Hamiltonian:

$$U = e^{-it \int d^3x \mathcal{H}}.$$

This picks out a time coordinate – doesn't manifestly respect Lorentz symmetry

- Lagrangian:

$$Z = \int \mathcal{D}\phi e^{i \int d^4x \mathcal{L}}$$

$$\mathcal{L}_{QCD} = \bar{\psi}(i\not{D})\psi - \frac{1}{4}(F_{\mu\nu}^i)^2 - m\bar{\psi}\psi$$

Approaches to QFT

- Perturbative – Taylor expand exponential. Asymptotic, not convergent. Requires small coupling
- Non-perturbative, exact – Only known for a few special cases, mostly in 2 spacetime dimensions
- Non-perturbative, numerical – Lattice. Gives static quantities (e.g., mass ratios), but not scattering amplitudes

No known efficient classical algorithm for scattering amplitudes, in general!

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Non-Relativistic vs. Relativistic Computation

Classically, believed that special relativity gives no change in computation power. (Aaronson)

- Same degrees of freedom – positions, momenta
- Time dilation – relativity has more power?
- Cosmic speed limit – relativity has less power?

Is the same true, once we add in QM?

- QFT is more fundamental, so it should be able to simulate non-relativistic quantum computers in polynomial time
- Is the converse true?

Difficulties in Simulating a QFT with a Quantum Computer

- Initial state does not determine final state – probabilistic
- Depending on incoming momentum, could have arbitrarily many output particles
- Field can take on infinitely many values – can we impose max field cutoff and discretize without introducing too much error?
- Infinitely many degrees of freedom – can we impose UV and IR cutoffs without introducing too much error?

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A Simple QFT – ϕ^4 theory

$$\mathcal{L} = \frac{1}{2}(\partial\phi)^2 - \frac{1}{2}m_0^2\phi^2 - \frac{\lambda_0}{4!}\phi^4$$

$$\mathcal{H} = \frac{1}{2}\pi^2 + \frac{1}{2}(\nabla\phi)^2 + \frac{1}{2}m_0^2\phi^2 + \frac{\lambda_0}{4!}\phi^4, \quad \pi = \dot{\phi}$$

Future study: gauge theories, higher-spin fields, zero mass.

Algorithm

Spatial dimension d . Hamiltonian – easier to use quantum simulation. Put theory on spatial lattice $\Omega = a\mathbb{Z}_{\hat{L}}^d$ (with periodic bc); $\mathcal{V} = \hat{L}^d$ lattice sites

- Input: List of incoming momenta
- Output: List of outgoing momenta. ϕ mass non-zero \Rightarrow # outgoing particles is at most linear in incoming particles' center-of-mass energy. Repeated runs sample from probability distribution for possible results of scattering experiment
- Precision: Probability of an outcome, according to simulation, differs from true probability by no more than $\pm\epsilon$

Representing the Field – Single Lattice Site

$n_b = O(\log(\phi_{\max}/\delta_\phi))$. How to choose ϕ_{\max}, δ_ϕ ? Note:
 $\pi_{\max} \sim \frac{1}{a^d \delta_\phi}$. Say in $|\psi\rangle$, where $\langle \psi | H | \psi \rangle \leq E$. Unlikely for
 $|\phi(x)|$ to be much larger than $O(\sqrt{E})$.

$$\phi_{\max} = O\left(\sqrt{\frac{\mathcal{V}E}{a^d m_0^2 \epsilon}}\right), \quad \pi_{\max} = O\left(\sqrt{\frac{\mathcal{V}E}{\epsilon a^d}}\right)$$
$$n_b = O\left(\log\left(\frac{\mathcal{V}E}{m_0 \epsilon}\right)\right)$$

Suzuki-Trotter

- 1 Prepare the free vacuum (Kitaev and Webb)
- 2 Excite wavepackets.
- 3 Adiabatically turn on interaction.
- 4 Simulate Hamiltonian time evolution.
- 5 Adiabatically turn off interaction.
- 6 Measure occupation numbers of momentum modes.

Excite wavepackets

$$\Gamma = \frac{2\pi}{\hat{L}a} \mathbb{Z}^d$$

$$\omega(\mathbf{p}) = \sqrt{m_0^2 + \frac{4}{a^2} \sum_{j=1}^d \sin^2\left(\frac{ap_j}{2}\right)}$$

$$a_{\mathbf{x}}^\dagger = \sum_{\mathbf{p} \in \Gamma} L^{-d} e^{-i\mathbf{p} \cdot \mathbf{x}} \sqrt{\frac{1}{2\omega(\mathbf{p})}} a_{\mathbf{p}}^\dagger$$

$$a_\psi^\dagger = \eta(\psi) \sum_{\mathbf{x} \in \Omega} a^d \psi(\mathbf{x}) a_{\mathbf{x}}^\dagger$$

$$H_\psi = a_\psi^\dagger \otimes |1\rangle\langle 0| + a_\psi \otimes |0\rangle\langle 1|$$

$$e^{-iH_\psi \pi/2} |0\rangle|0\rangle = -i a_\psi^\dagger |0\rangle|1\rangle$$

Adiabatically turn on interaction – How slowly do we go?

- Rate at which we turn on λ_0 depends on the physical mass.
- Weak coupling: can calculate perturbatively
- Strong coupling: slowly increase λ_0 , estimating the mass at each value. Measure energy using phase estimation
- Route in parameter space may be circuitous – avoid “phase transition” where mass goes to 0

Representing the Field – IR Cutoff

Want V infinite – non-interacting, at first. Non-zero mass, so errors shrink exponentially with distance δ :

$\epsilon \sim e^{-\delta/m} \Rightarrow V \sim n_{\text{in}} \log(1/\epsilon)$. Similarly, non-interacting at end, so $V \sim n_{\text{out}} \log(1/\epsilon)$. $p \ll 1/a$, so

$$\mathcal{V} \sim n_{\text{out}}^{d+1}$$

Representing the Field – UV Cutoff

$$\mathcal{L}_{\text{eff}} = \mathcal{L}^{(0)} + \frac{c}{6!}\phi^6 + c'\phi^3\partial^2\phi + \frac{c''}{8!}\phi^8 + \dots$$

Dimensional analysis: later terms get larger power of a

Performance: Strong Coupling, $d = 1, 2$

	$\lambda_c - \lambda_0$	p	n_{out}
$d = 1$	$\left(\frac{1}{\lambda_c - \lambda_0}\right)^{8+o(1)}$	$p^{4+o(1)}$	$\tilde{O}(n_{\text{out}}^5)$
$d = 2$	$\left(\frac{1}{\lambda_c - \lambda_0}\right)^{5.04+o(1)}$	$p^{6+o(1)}$	$\tilde{O}(n_{\text{out}}^{7.128})$

Table: $f(n) = \tilde{O}(g(n))$ means $f(n) = O(g(n) \log^c(n))$ for some c

Performance: Weak Coupling

$$G_{\text{weak}} \sim \begin{cases} \left(\frac{1}{\epsilon}\right)^{1.5+o(1)} & : d = 1 \\ \left(\frac{1}{\epsilon}\right)^{2.376+o(1)} & : d = 2 \\ \left(\frac{1}{\epsilon}\right)^{3.564+o(1)} & : d = 3 \end{cases}$$